

UT Environmental Science Institute

#### #18

## Supermassive Black Holes: Galaxy Monsters

### Dr. John Kormendy April 19, 2002

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## **Supermassive Black Holes in Galactic Nuclei**

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## April 19, 2002

# Est Environmental Science Institute

What are black holes? Why are we looking for them? The search for supermassive black holes Supermassive black holes and galaxy formation

## **Black Hole**

Suppose we squeeze a planet of fixed mass *M* into smaller and smaller volumes.

The surface gravity gets bigger as you make the planet smaller. Surface gravity is proportional to 1/radius<sup>2</sup>.



## To turn the Earth into a black hole, we would have to squeeze it into the size of a grape.



## **Two Types of Black Holes**

Black holes with masses of a few Suns are well understood.

The most massive stars turn into such black holes when they die in supernova explosions.









These are millions of stellar-mass black holes — dead stars — scattered throughout the disk of our Milky Way galaxy.

## Sombrero Galaxy: Black Hole Mass = 1 Billion Suns

## **Two Types of Black Holes**

Supermassive black holes with masses of a million to a few billion Suns live in galactic centers.

We understand some of what they do, but we don't know where they come from.



### **The Discovery of Quasars**

The identification by Maarten Schmidt (1963, Nature, 197, 1040) of the radio source 3C 273 as a 13<sup>th</sup> magnitude "star" with a redshift of 16 % of the speed of light was a huge shock. The Hubble law of the expansion of the Universe implies that 3C 273 is one of the most distant objects known. It must be enormously luminous.

Quasars — violently active galactic nuclei — are the most luminous objects in the Universe, more luminous than any galaxy.

The energy requirements for powering quasars were the first compelling argument for black hole engines.





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# Many radio galaxies and quasars have jets that feed lobes of radio emission



Cygnus A

## **Supermassive Black Holes as Quasar Engines**

Let's try to explain quasars using nuclear reactions like those that power stars:

• The total energy output of a quasar is at least the energy stored in its radio halo  $\approx 10^{54}$  Joule.

• Via  $E = mc^2$ , this energy "weighs" 10 million Suns.

- But nuclear reactions have an efficiency of only 1 %.
- So the waste mass left behind in powering a quasar is 10 million Suns / 1  $\% \approx$  1 billion Suns.
- Rapid brightness variations show that a typical quasar is no bigger than our Solar System.
- But the gravitational energy of 1 billion Suns compressed inside the Solar System  $\approx 10^{55}$  Joule.

"Evidently, although our aim was to produce a model based on nuclear fuel, we have ended up with a model which has produced more than enough energy by gravitational contraction. The nuclear fuel has ended as an irrelevance."

Donald Lynden-Bell (1969)

This argument convinced many people that quasar engines are supermassive black holes that swallow surrounding gas and stars.





## Why Jets Imply Black Holes — 1



Jets remember ejection directions for a long time. This argues against energy sources based on many objects (supernovae). It suggests that the engines are rotating gyroscopes - rotating black holes.

## Why Jets Imply Black Holes — 2



Jet knots move at almost the speed of light. This implies that their engines are as small as black holes. This is the cleanest evidence that quasar engines are black holes.

## Why Jets Imply Black Holes — 2

Superluminal Motion in the M87 Jet



Jet knots in M87 look like they are moving at 6 times the speed of light (24 light years in 4 years).

This means that they really move at more than 98 % of the speed of light.

## **Supermassive Black Holes as Quasar Engines**

The huge luminosities and tiny sizes of quasars can be understood if they are powered by black holes with masses of a million to a few billion Suns.

Gas near the black hole settles into a hot disk, releasing gravitational energy as it spirals into the hole.

Magnetic fields eject jets along the black hole rotation axis.

# A black hole lights up as a quasar when it is fed gas and stars.

## How do you feed a quasar?

Simulation: Josh Barnes

One possible answer: Galaxy collisions and mergers dump gas into the center.

## PROBLEM

People believe the black hole picture. They have done an enormous amount of work based on it.

But for many years there was no direct evidence that supermassive black holes exist.

So the search for supermassive black holes became a very hot subject.

Danger: It is easy to believe that we have proved what we expect to find. So the standard of proof is very high.

## The Quasar Era Was More Than 10 Billion Years Ago



## The Search For Supermassive Black Holes

The image shown here in the lecture is not available for public distribution.

Canada-France-Hawaii-Telescope



## Simple Spectrograph





M 31 on spectrograph slit

Spectrum of M 31 The brightness variation of the galaxy has been divided out. The zigzag in the lines is the signature of the rapidly rotating nucleus and central black hole.





Distance  $\approx 2.5 \times 10^6$  light years (ly)  $\Rightarrow$  scale = 82 arcsec / 1000 ly.

Observe rotation velocity V  $\approx$  160 km/s at radius r = 1 arcsec.

Mass: Balance centripetal acceleration and gravity:

$$\begin{array}{ll} mV^2/r = GMm/r^2 & \Rightarrow & M = V^2 \, r \, / \, G \, , \\ & \text{where } M = \text{central ("black hole") mass;} \\ & m = \text{mass of orbiting star;} \\ & G = Gravitational \ constant; \\ & = 6.673 \, x \, 10^{-8} \, \text{cm}^3 \, / \, \text{g} \, \text{s}^2. \end{array}$$

#### **Approximate mass**

 $(160 \text{ km/s})^2 (10^5 \text{ cm / km})^2 (1 \text{ arcsec}) (9.47 \text{ x } 10^{17} \text{ cm/ly})$ 

V<sup>2</sup> r / G ≈ \_\_\_\_\_

 $(6.673 \times 10^{-8} \text{ cm}^3 \text{ g}^{-1} \text{ s}^{-1}) (1.99 \times 10^{33} \text{ g/ M}_{\odot}) (82 \text{ arcsec / } 1000 \text{ ly})$ 

= 22 x  $10^6 M_{\odot}$ . This is an underestimate because it does not include the effects of projection or random velocities.

## **NGC 3115: Black Hole Mass = 1 Billion Suns**



## **NGC 3115: Black Hole Mass = 1 Billion Suns**



NGC 3115 has a bright central cusp of stars like we expect around a black hole.

Stars in this nuclear cluster move at about 1000 km/s.

But: if the nucleus contained only stars and not a black hole, then its escape velocity would be 350 km/s. Stars moving at 1000 km/s would fly away.

This shows that the nucleus contains a dark object of mass 1 billion Suns.



## **The Nuker Team**



Doug Richstone





Karl Gebhardt

Sandra Faber



John Kormendy



Alan Dressler



Alex Filippenko



**Ralf Bender** 



**Richard Green** 

Thanks Also To STIS Team Members Gary Bower Mary Beth Kaiser Charlie Nelson

## Black Hole Census — April 2002

Galaxy	Distance B (million ly)	lack Hole Mass (million Suns)	Galaxy	Distance E (million ly)	Black Hole Mass (million Suns)	
Milky Way	0.028	3	NGC 5128	6.5	200	
M 31	2.3	40	NGC 2787	42	71	
M 32	2.3	3	M 87	52	2500	
M 81	12.7	68	NGC 4350	55	600	
NGC 3115	32	1000	NGC 4459	55	73	
NGC 4594	32	1000	NGC 4596	55	78	
NGC 3379	34	100	NGC 4374	60	1000	
NGC 3377	37	100	IC 1459	95	200	
NGC 1023	37	39	NGC 4261	104	540	
NGC 3384	38	14	NGC 7052	192	330	
NGC 4697	38	120	NGC 6251	345	600	
NGC 7457	43	3				
NGC 4564	49	57	NGC 4945	12.1	1	
NGC 4342	50	300	NGC 4258	23	42	
NGC 4486B	50	500	NGC 1068	49	17	
NGC 4742	51	14				
NGC 4473	51	100				
NGC 4649	55	2000	Kormendy et al.			
NGC 2778	75	20	Gebhardt et nuk.			
NGC 3608	75	110	STIS GTO Team			
NGC 7332	75	15				
NGC 821	78	50				
NGC 4291	85	250				
NGC 5845	85	320				

## M 87: Black Hole Mass = 2.5 Billion Suns





M 87 was observed with Hubble by Harms, Ford, and collaborators.

#### Spectrum of Gas Disk in Active Galaxy M87



From the difference in Doppler shifts seen on opposite sides of the center, the disk rotates at almost 600 km/s. This implies a black hole of mass 2.5 billion Suns.

## M 84: Black Hole Mass = 1 Billion Suns

The Space Telescope Imaging Spectrograph is providing spectacular new data on black holes.



Hubble Space Telescope

Bower et al. 1998, ApJ, 492, L111

# We observe only a shadow of the dark object — its gravitational effect on ordinary stars and gas.









## **Could we be detecting a cluster of dark stars?**



#### **Possibilities**

Brown dwarf stars White dwarf stars Neutron stars Stellar-mass black holes

## NGC 4258: Dark Mass = 42 Million Suns

## Our Galaxy: Dark Mass = 3 Million Suns

We can measure the rotation of these two galaxies especially close to the center.

The dark mass is in such a small volume at the center that alternatives to a black hole (failed stars or dead stars) are ruled out.

These are the best black hole candidates.

### Have we discovered black holes in galactic nuclei?

Probably.

Other alternatives are very implausible.

But: Absolute proof requires that we see velocities of almost the speed of light from near the surface of the black hole.



X-ray observations of Seyfert galaxies show spectral lines as wide as 100,000 km/s.

This is 1/3 of the speed of light.

The image shown here in the lecture is not available for public distribution.

## The bulgeless galaxy M 33 does not contain a black hole.



Typical stars in the nucleus of M 33 move at only 21 km/s. Any black hole must be less massive than 1500 Suns.

## Conclude: Every galaxy that contains a bulge component also contains a black hole.

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## **Black Hole Conclusions**

# Black hole masses are just right to explain the energy output of quasars.



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Bigger black holes live in bigger galaxy bulges.

Bigger black holes live in bulges in which the stars move faster.



These diagrams contain almost the same information, because galaxy mass is measured by stellar velocity. What is new in the right panel? **Answer: how much the bulge collapsed when it formed.** 



Black hole mass is more connected with how much a bulge collapsed than with how big the bulge is. So there is a close connection between black hole growth and bulge formation. CONCLUSION The formation of bulges and the growth of their black holes, when they shone like quasars, happened together.

This unifies two major areas of extragalactic research: quasars and galaxy formation.





#### Short-term

#### It should be possible to derive black hole masses for thousands of quasars and other active galaxies and probe the growth of black holes as the Universe evolved.

Long-term

**Gravitational wave astronomy** 

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Dr. John Kormendy is a professor at the University of Texas at Austin. His research interests include the search for black holes in galactic nuclei, elliptical galaxies and bulges of disk galaxies, dark matter, and secular evolution in galaxy structure.